A classical painting depicting a winged angel standing triumphantly on a fallen figure. The angel, with large, feathered wings and a flowing robe, holds a sword aloft in his right hand and a scroll in his left. The fallen figure lies on the ground, head down, with a sword resting near his hand. The background is a soft, hazy landscape. The overall tone is dramatic and heroic.

The next ten years: chaos or illumination

Pier Oddone, Fermilab
Brookhaven Forum 2010
May 28, 2010

Hold on to your seats.....

- The next decade is here! The greatest expectations in generations and the greatest uncertainties....
- As we have heard in BF2010: any number of models are possible – few clues from present experiments.
- The LHC will be the dominant tool – but critical questions remain in other areas: for neutrinos, rare processes, dark matter and dark energy

New physics: what clues do we have?

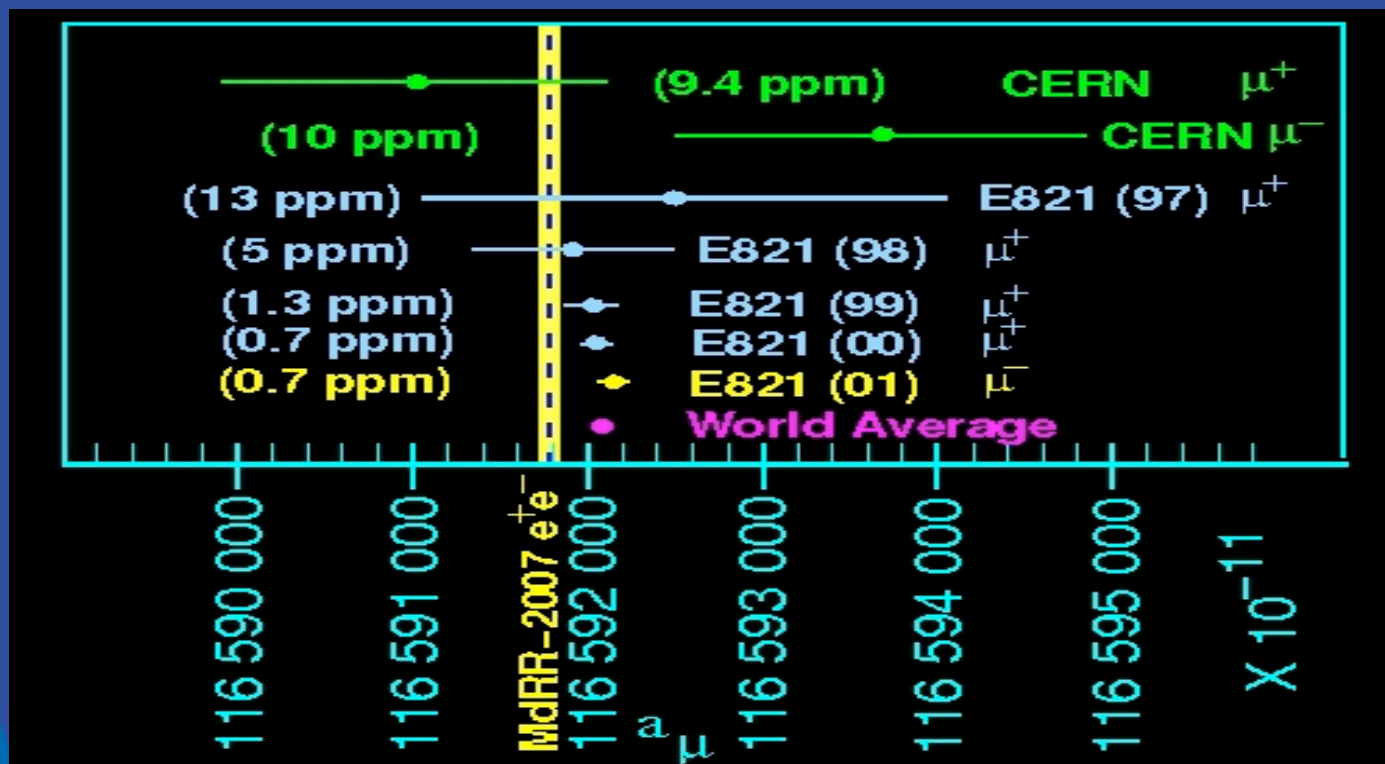
- The biggies:
 - Dark matter
 - Dark energy
 - Neutrino masses
 - Matter antimatter asymmetry in the universe
 - Three flavors
 - Unitarity at TeV scale?
- But is the new physics around the corner?
Looking through precision tests.....

What clues do we have?

- G-2: anomalous magnetic moment of the muon shows a 3 sigma deviation from SM predictions



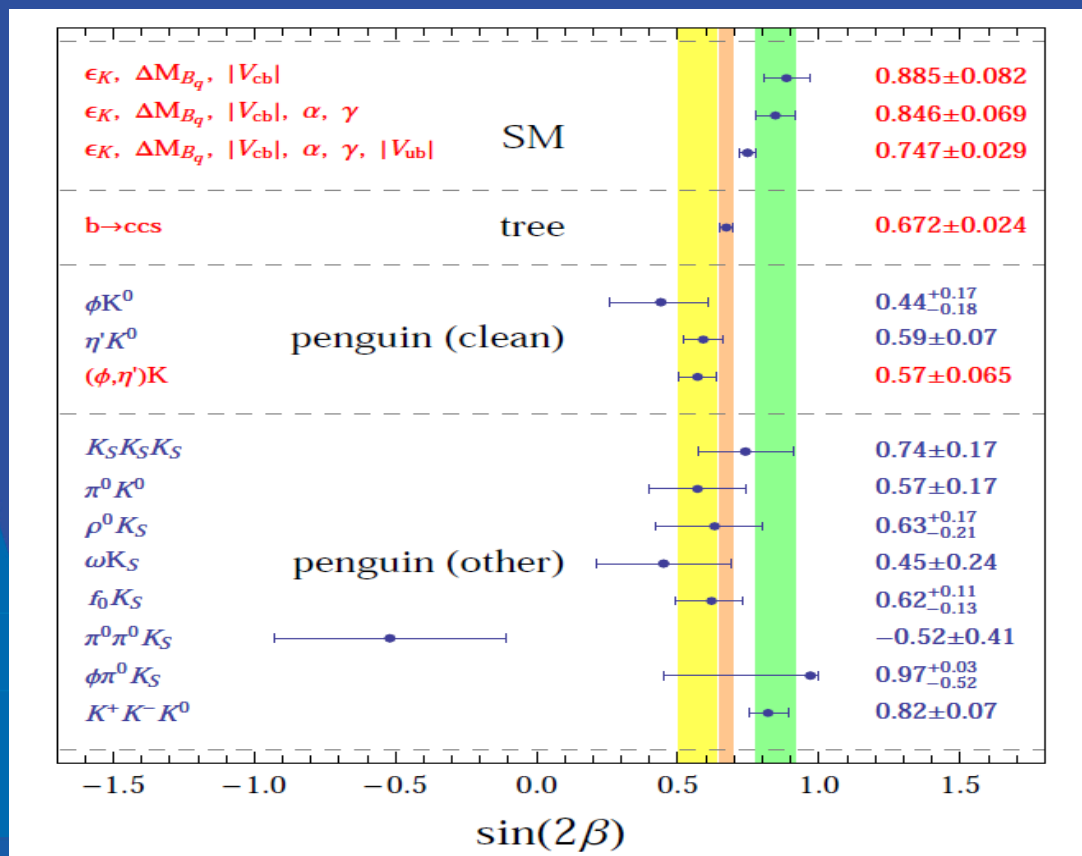
What clues do we have?



$$\Delta a_\mu \equiv a_\mu^{exp} - a_\mu^{SM} = (255 \pm 80) \times 10^{-11}$$

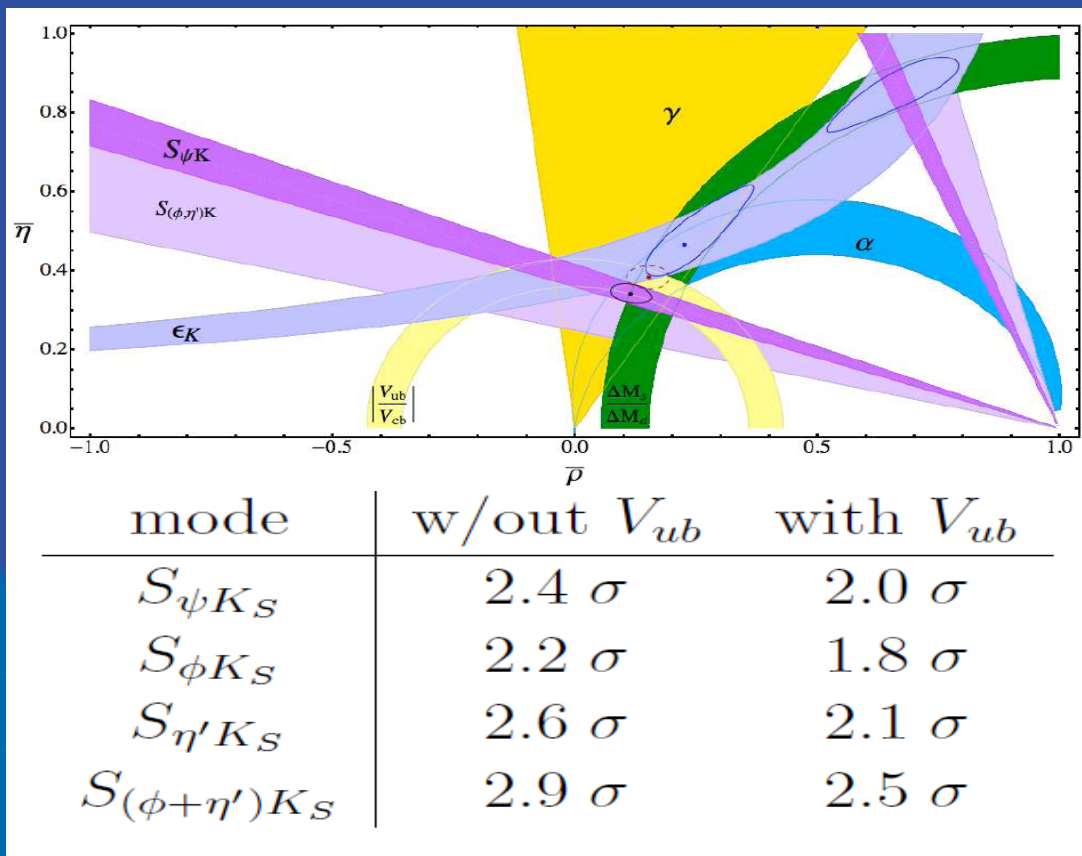
What clues do we have?

Tension in comparison of different determinations of $\sin(2\beta)$ (See Neubert at BF2010)



What clues do we have?

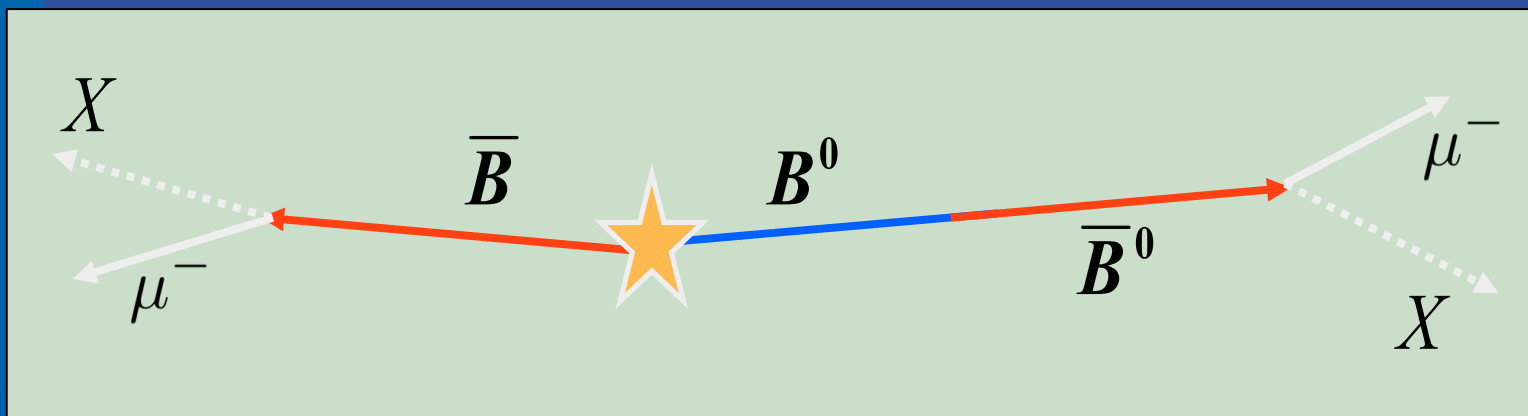
Tensions in fitting unitarity triangle



Lunghi and Soni, PL B666 (2008) 162

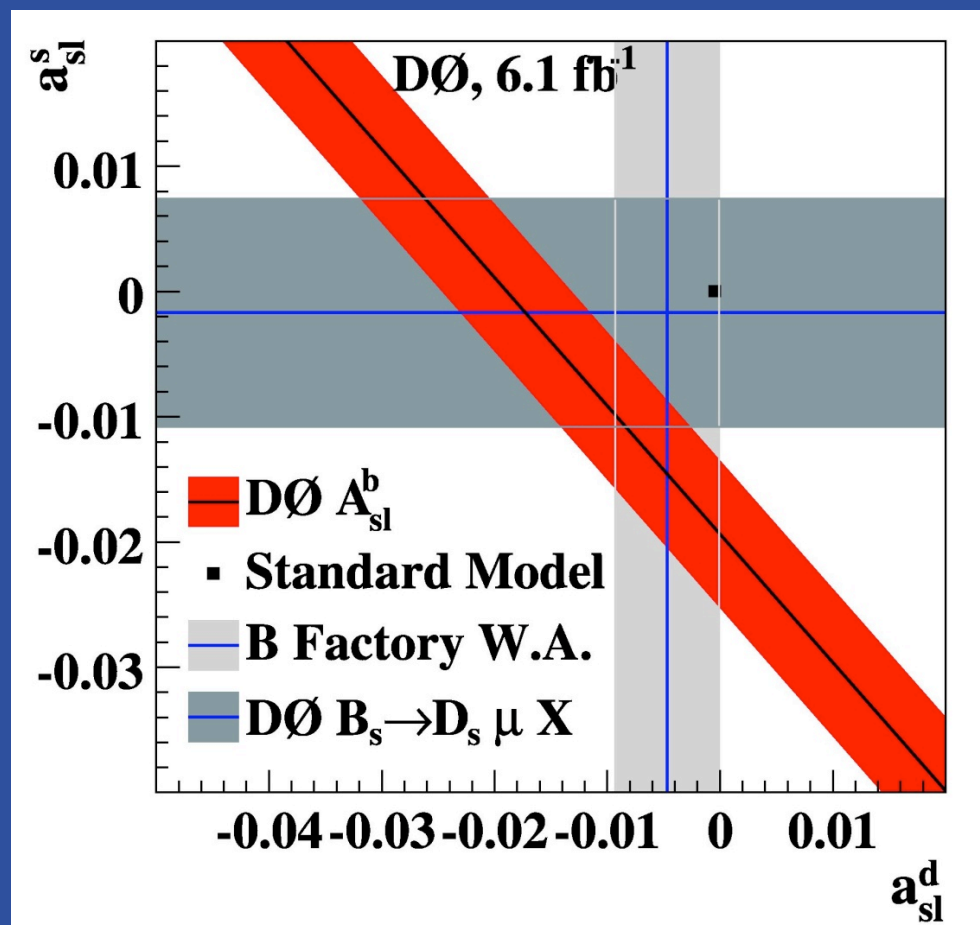
What clues do we have?

- Recent measurement of di-muon asymmetry in DZERO



$$A_{sl}^b \equiv \frac{N_b^{++} - N_b^{--}}{N_b^{++} + N_b^{--}}$$

What clues do we have?

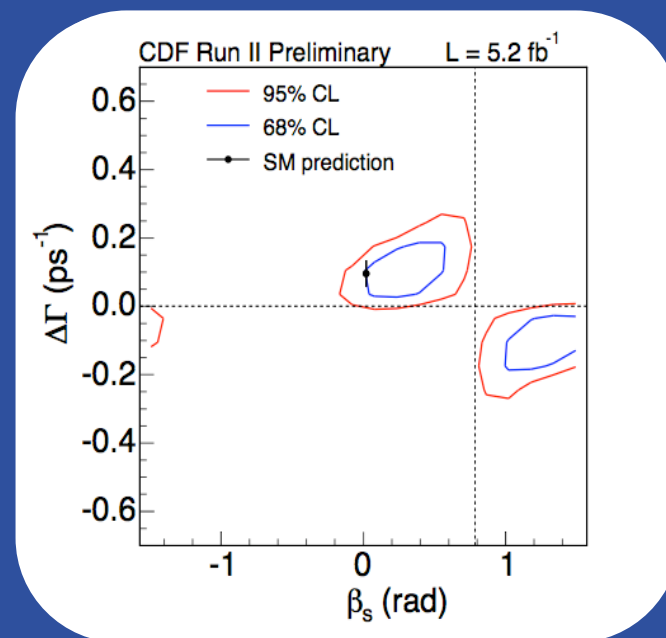
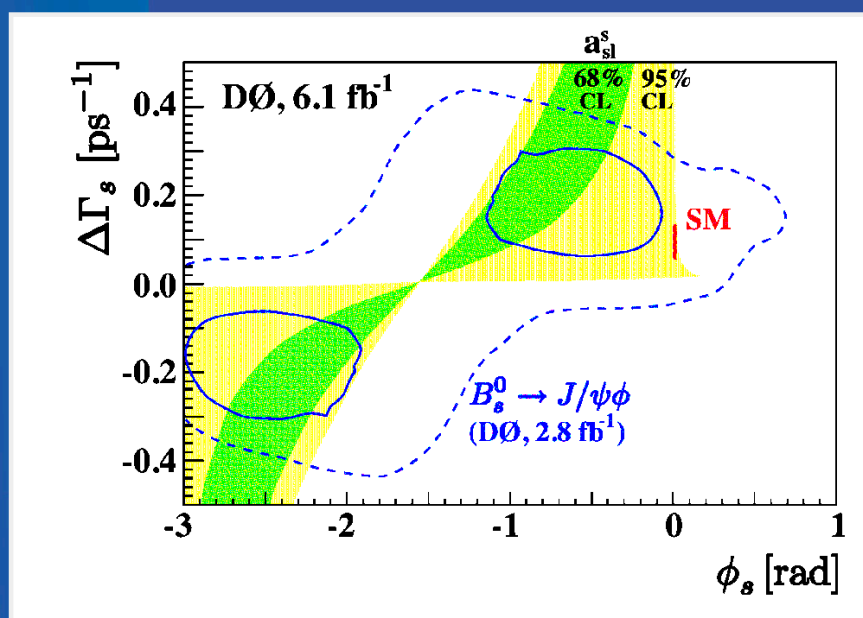


New at
BF2010,
Tonelli,
Tsybychev

$$A_{sl}^b = 0.506 a_{sl}^d + 0.494 a_{sl}^s$$

What clues do we have?

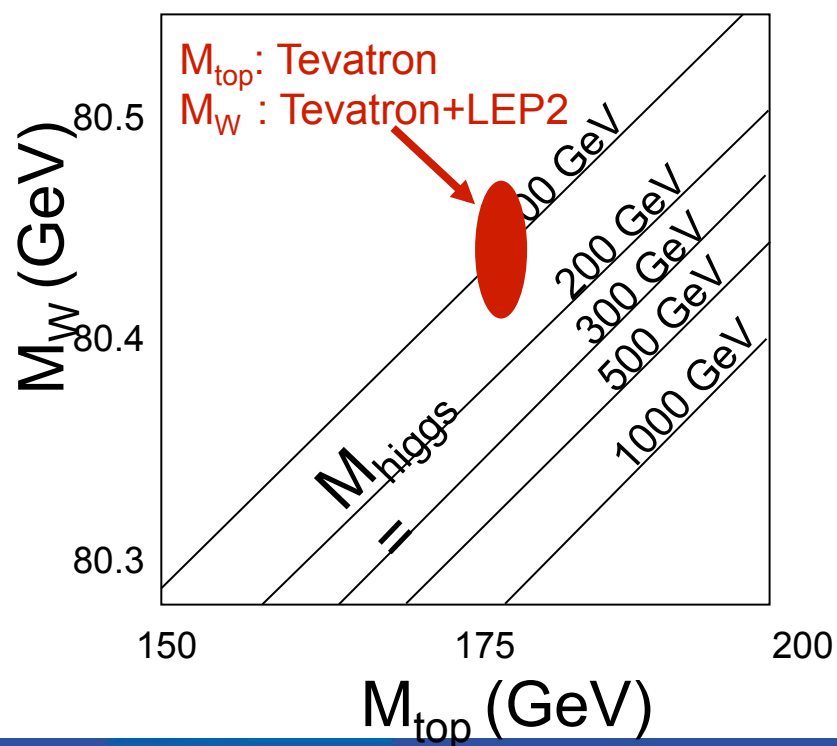
Further evidence from independent measurements of ϕ_s and $\Delta\Gamma_s$ in $B_s \rightarrow J/\psi\phi$ decay (Tonelli, Tsybychev @BF2010)



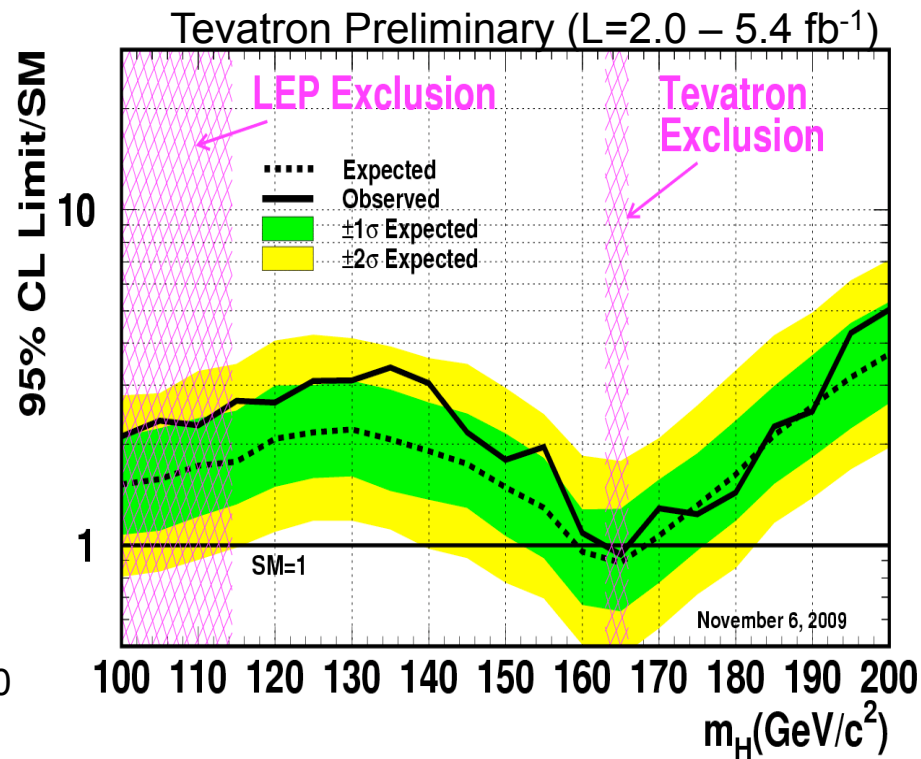
What clues do we have?

Electroweak fits

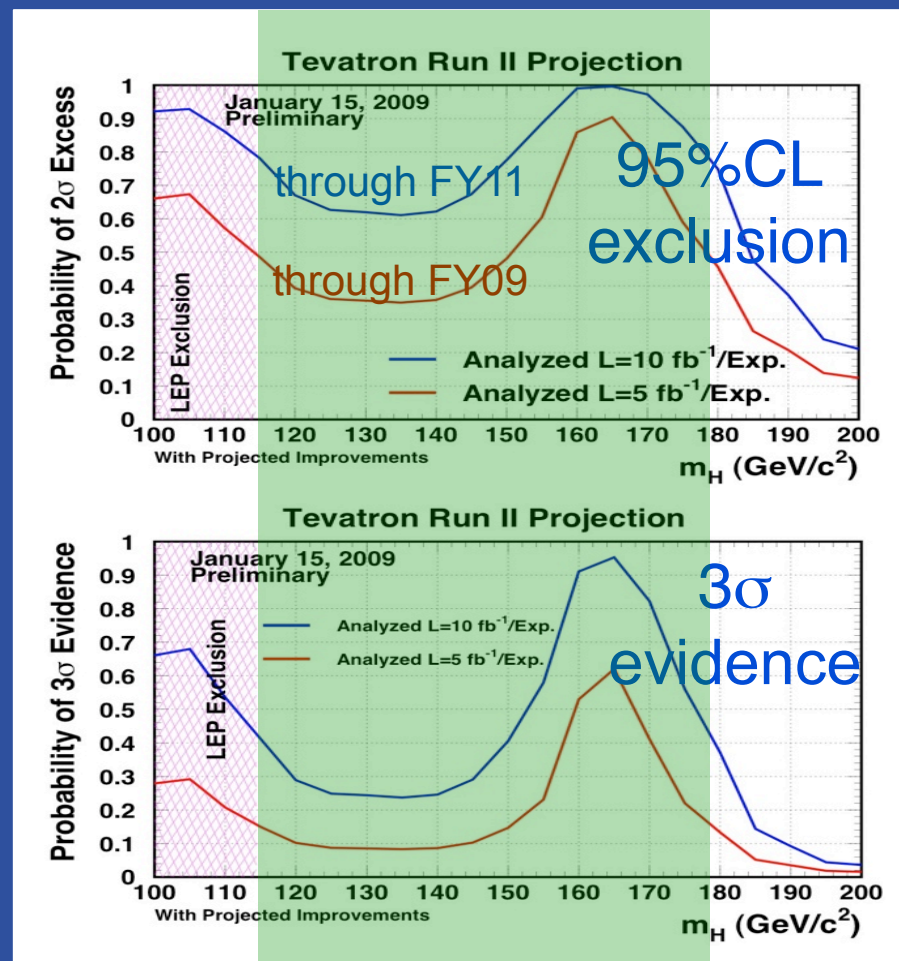
Predict Higgs Mass



Search for Higgs



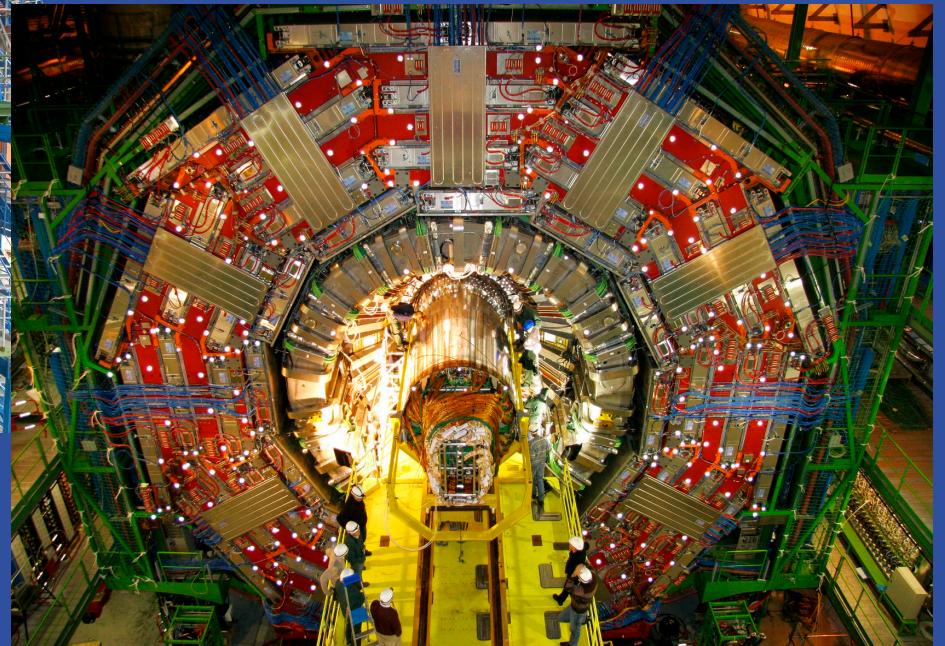
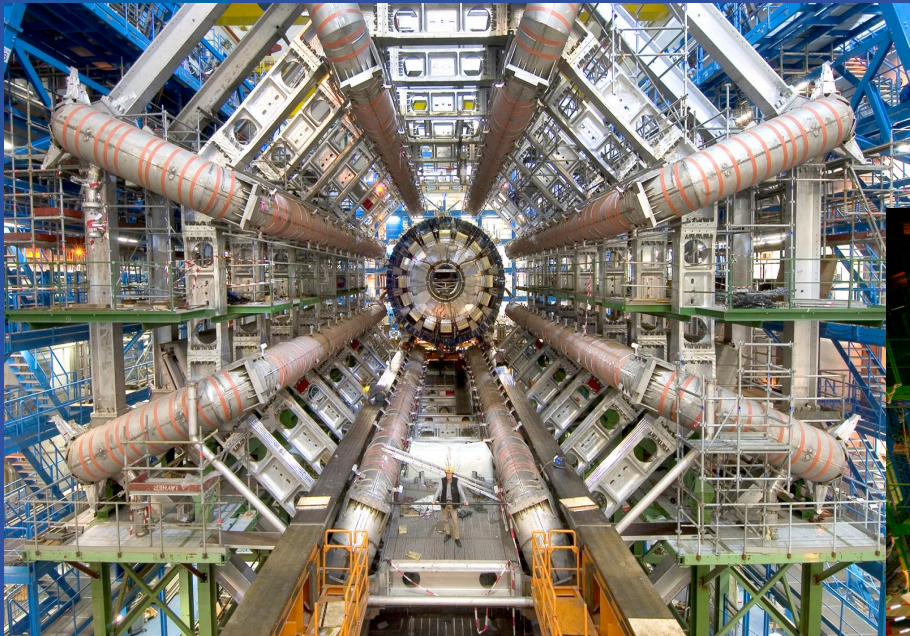
Is there more juice in the Tevatron?



Energy frontier will move to the LHC



Powerful detectors

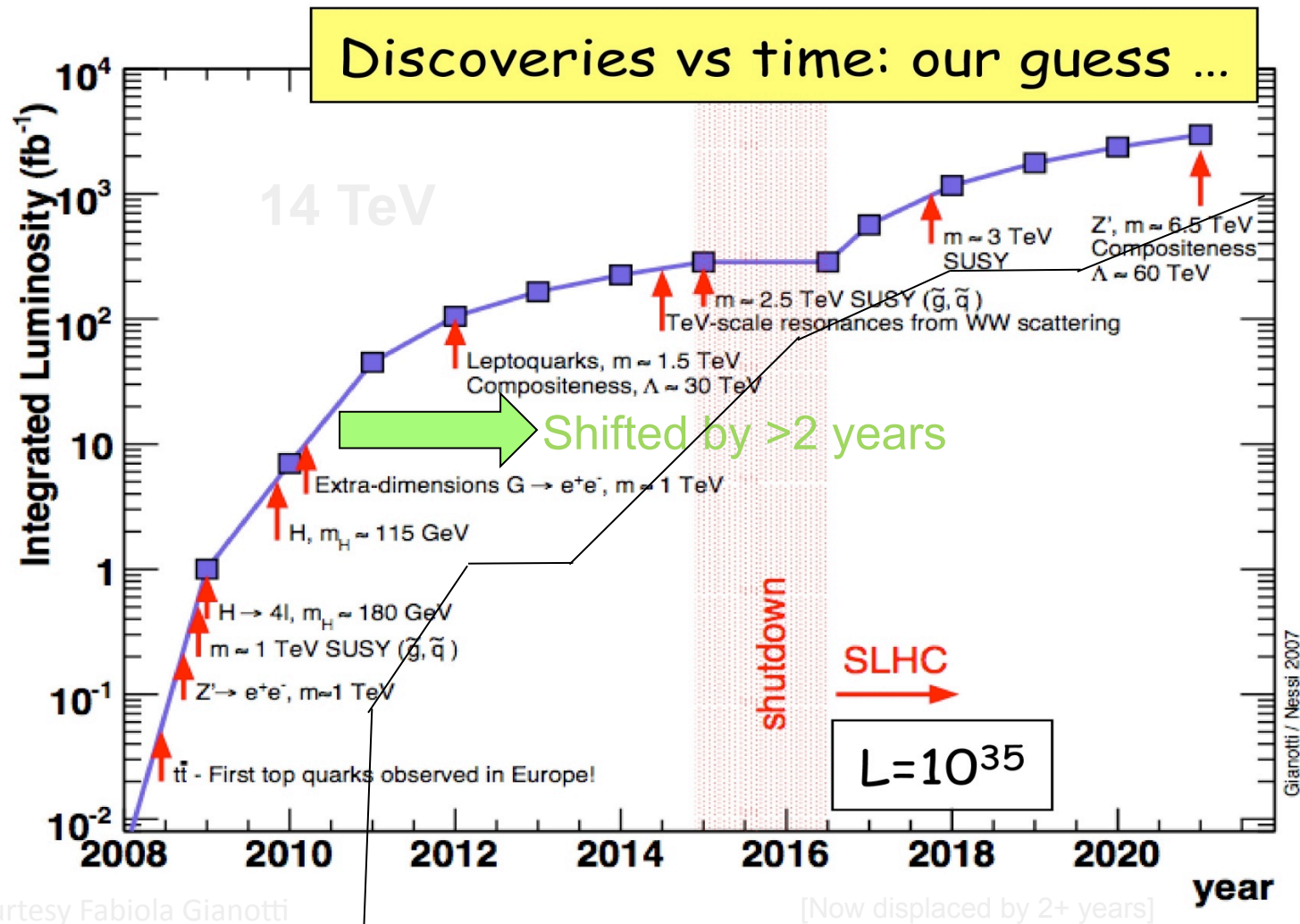


(See Erbacher and
Heinemann at BF2010)

Huge physics reach

- Theorists have imagined a veritable Jurassic Park of possibilities in this new energy range:
 - Minimal Supersymmetric (MSSM)
 - Many other models of supersymmetry (less predictive)
 - New Z' similar to the Z boson, but higher masses
 - Extra dimensions and Kalusa-Klein towers
 - Technicolor
 - Lepto-quarks
 - Mini black holes
 -

LHC physics reach (3 years ago)



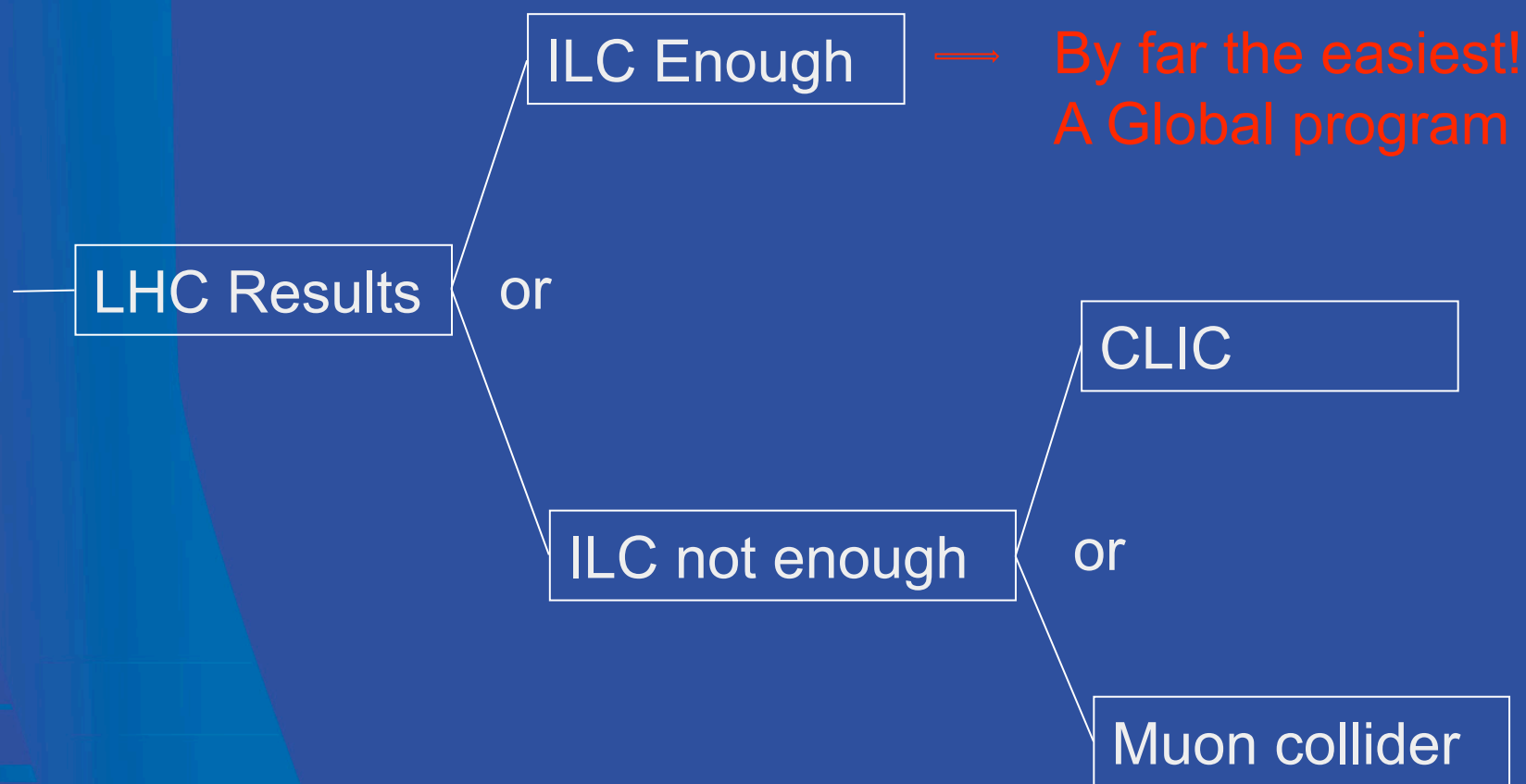
Courtesy Fabiola Gianotti

[Now displaced by 2+ years]

Detectors are fantastic!

- Physics will come quickly: weeks to months after data taking
- Also means many effects will appear first at the 3 sigma level so we will be chasing many things on the way to discovery
- When will we be able to be confident enough to say what the next machine is?
- Great confusion? (See Lykken at BF2010)

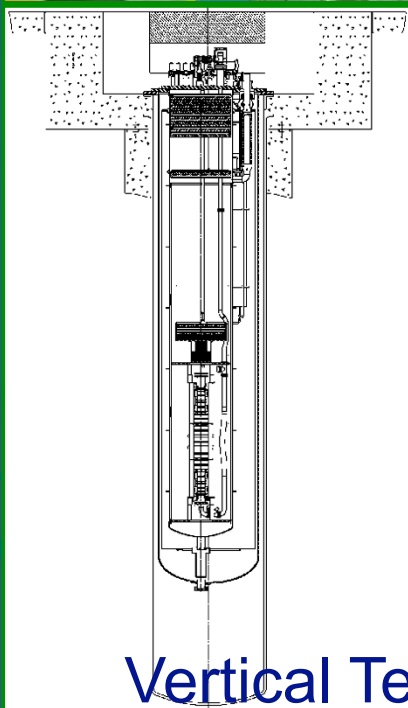
Biggest decision of the decade !



ILC/Project X/XFEL technology



Horizontal Test Stand



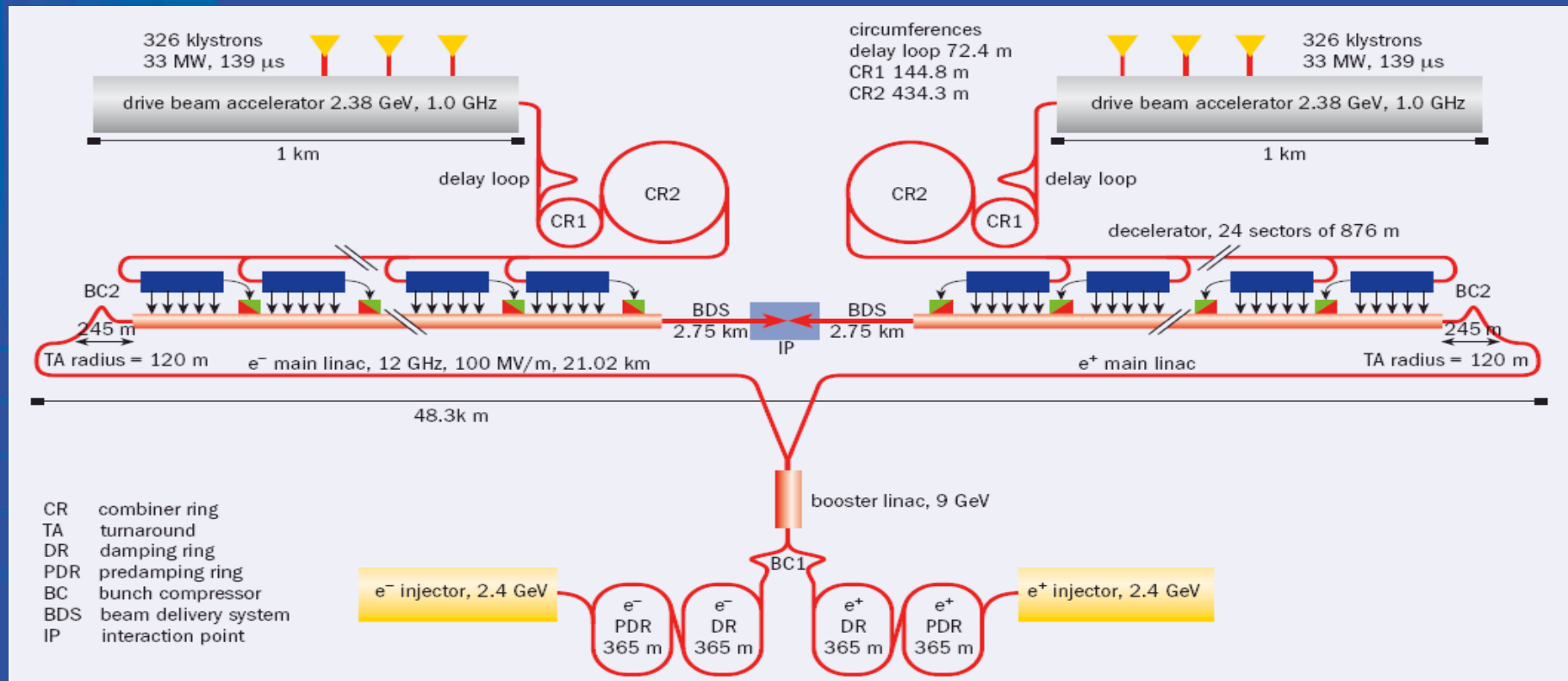
Vertical Test Stand



1st cryomodule

If we need higher energies.....

- If the energy of the ILC is too small (0.5 TeV or a little higher) we will need another approach: CLIC or muon collider

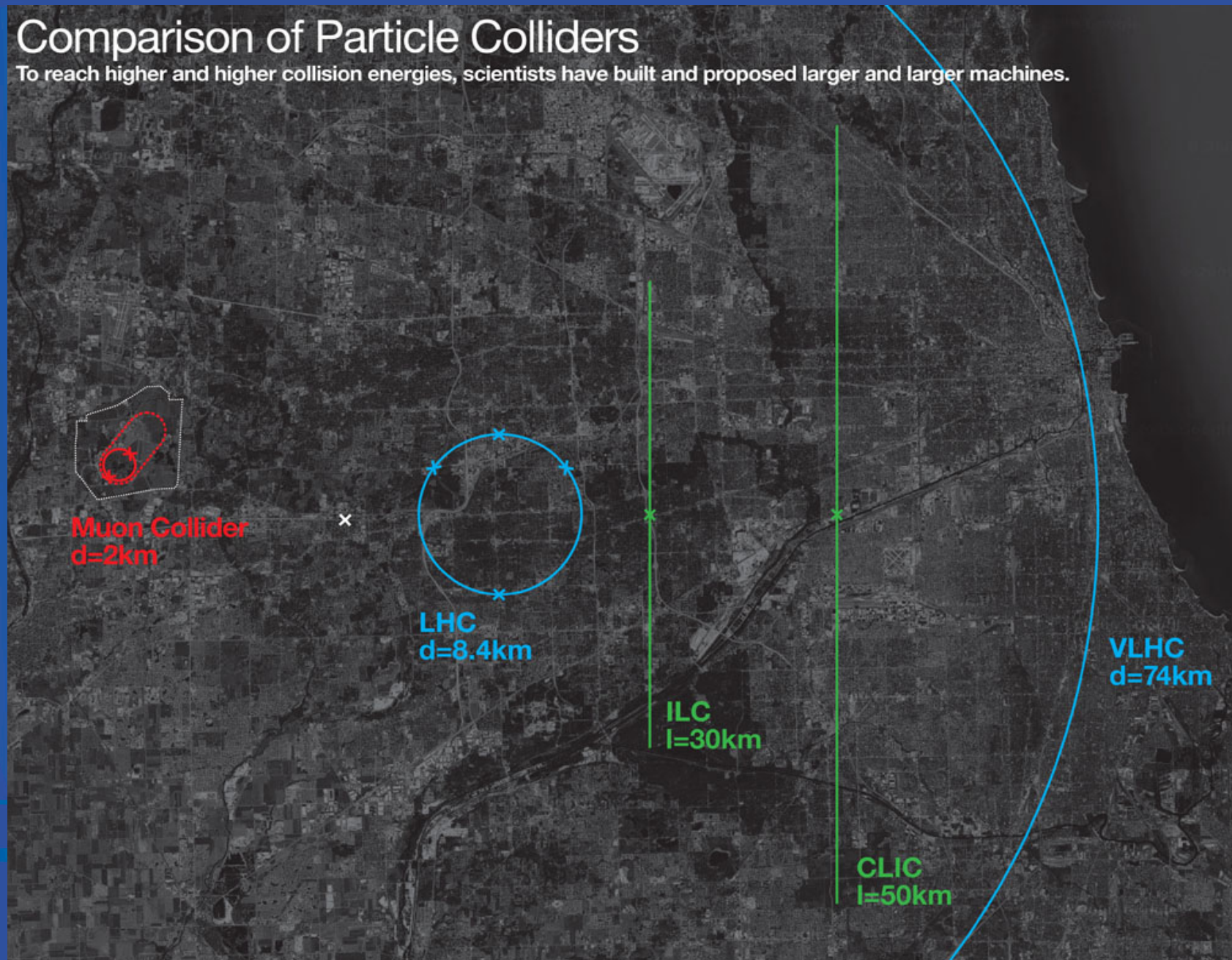


Muon Collider approach

- If the energy of the ILC is too small (0.5 TeV or a little higher) we will need another approach: CLIC or muon collider
- Collider based on a secondary beam: we have experience basing colliders on antiprotons. For muons we must do it in 20 msec.
- The biggest advantages are: narrow energy spread (no beamstrahlung) and small physical footprint (no synchrotron radiation)
- DOE OHEP has asked Fermilab to organized the national R&D program

Comparison of Particle Colliders

To reach higher and higher collision energies, scientists have built and proposed larger and larger machines.



Muon Collider Conceptual Layout

Project X

Accelerate hydrogen ions to 8 GeV using SRF technology.

Compressor Ring

Reduce size of beam.

Target

Collisions lead to muons with energy of about 200 MeV.

Muon Cooling

Reduce the transverse motion of the muons and create a tight beam.

Initial Acceleration

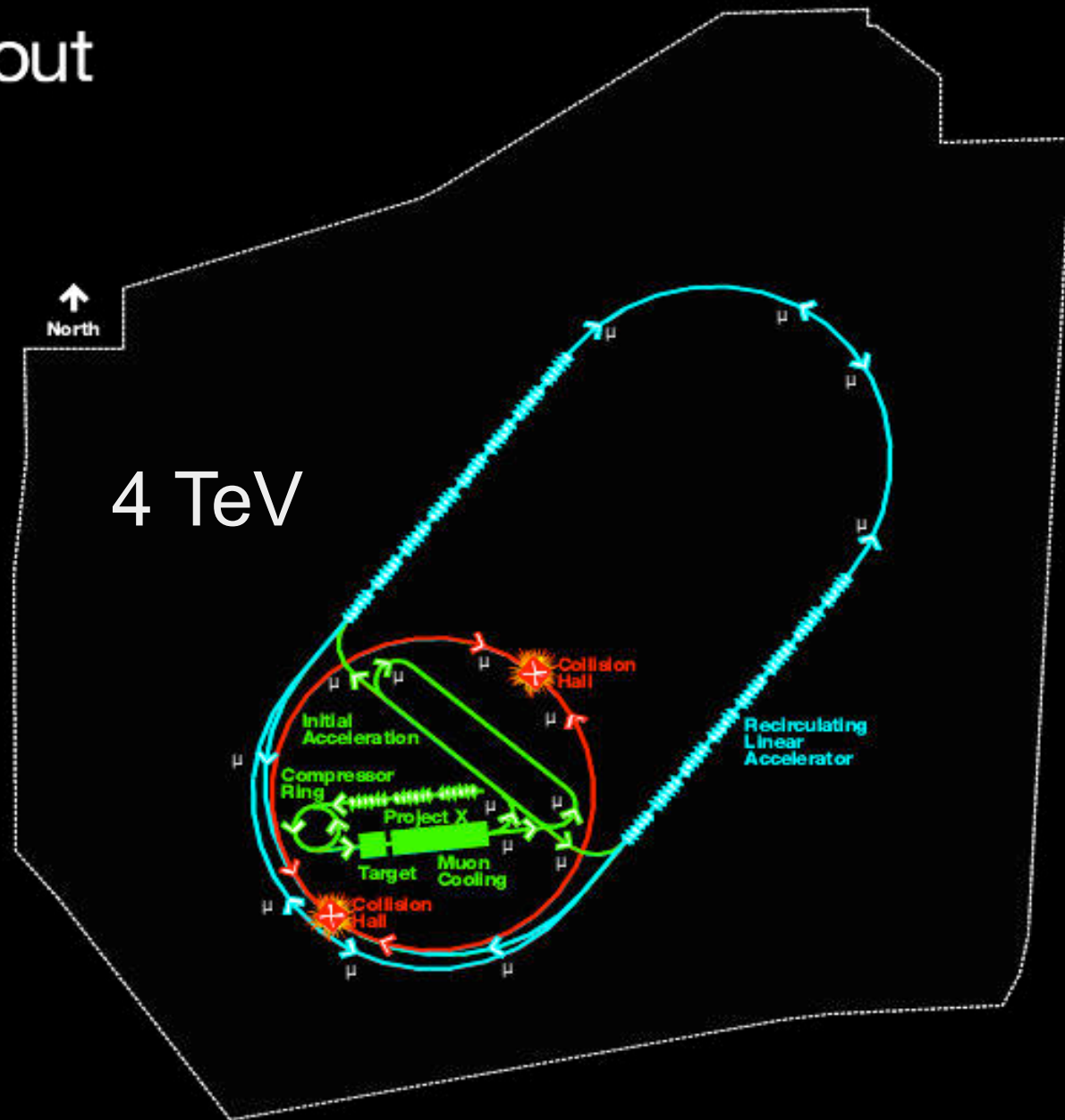
In a dozen turns, accelerate muons to 20 GeV.

Recirculating Linear Accelerator

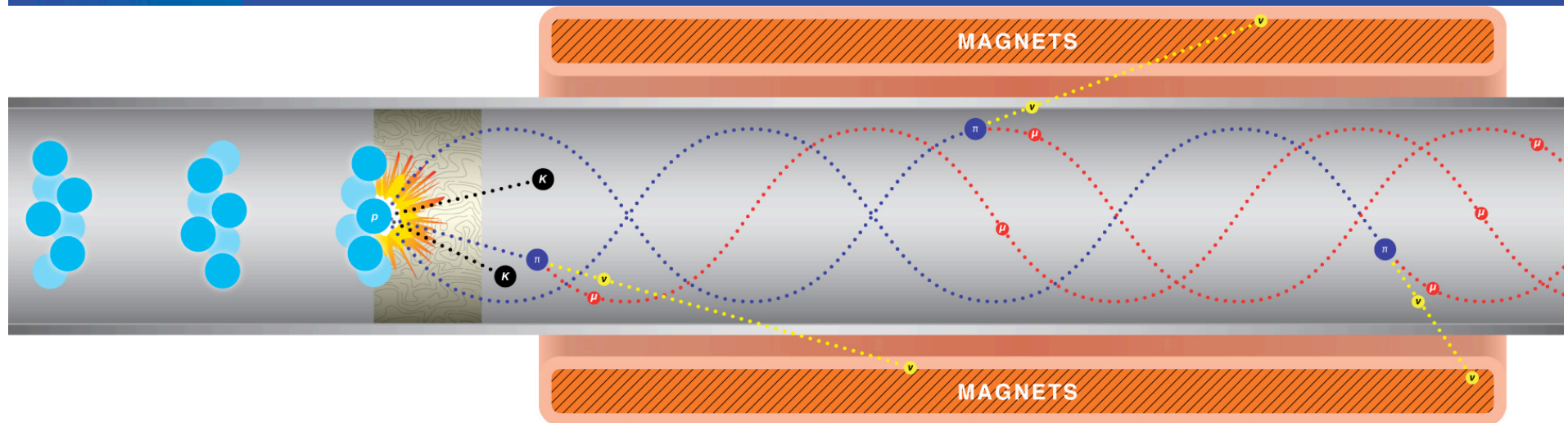
In a number of turns, accelerate muons up to 2 TeV using SRF technology.

Collider Ring

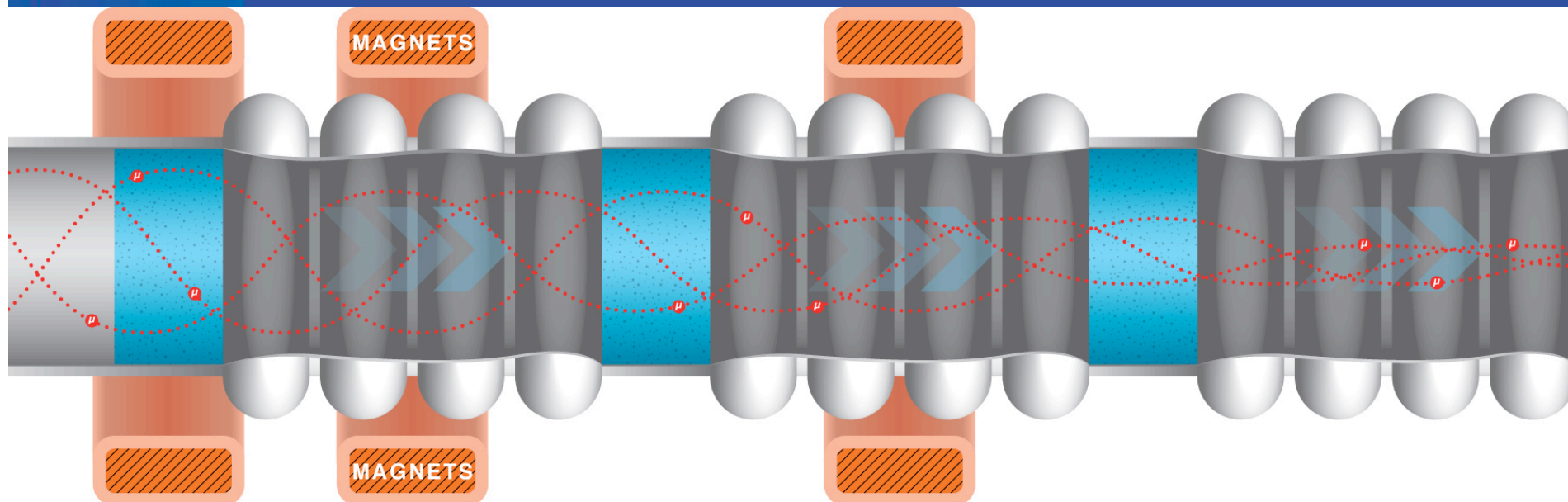
Located 100 meters underground. Muons live long enough to make about 1000 turns.



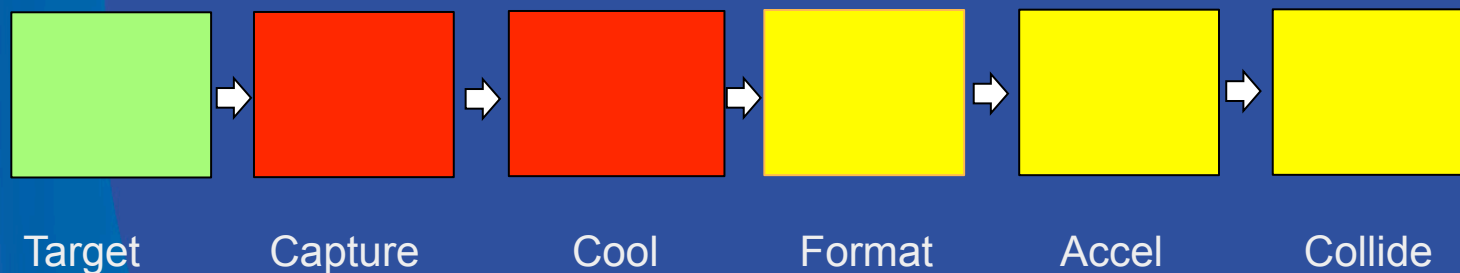
Targeting and capturing



Capturing and cooling

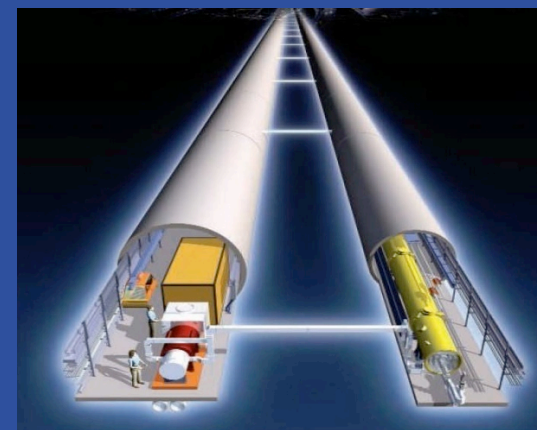
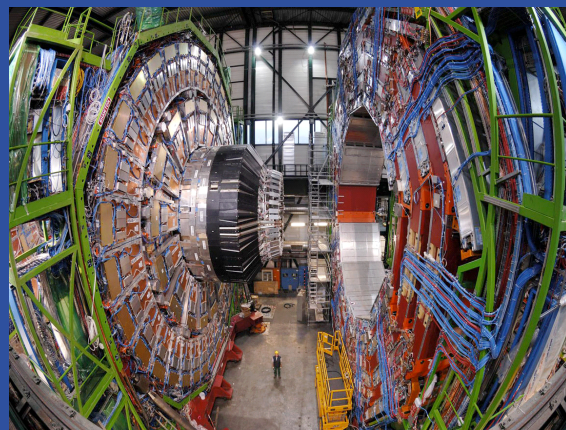
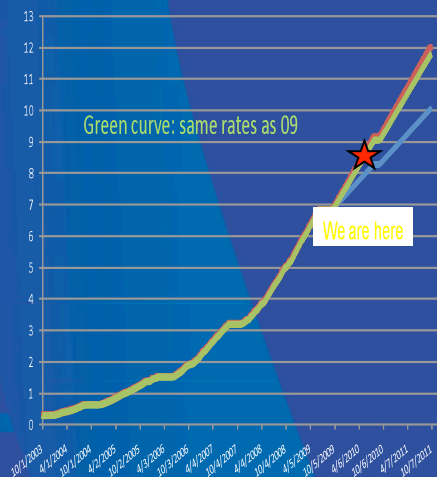


Muon collider functional layout

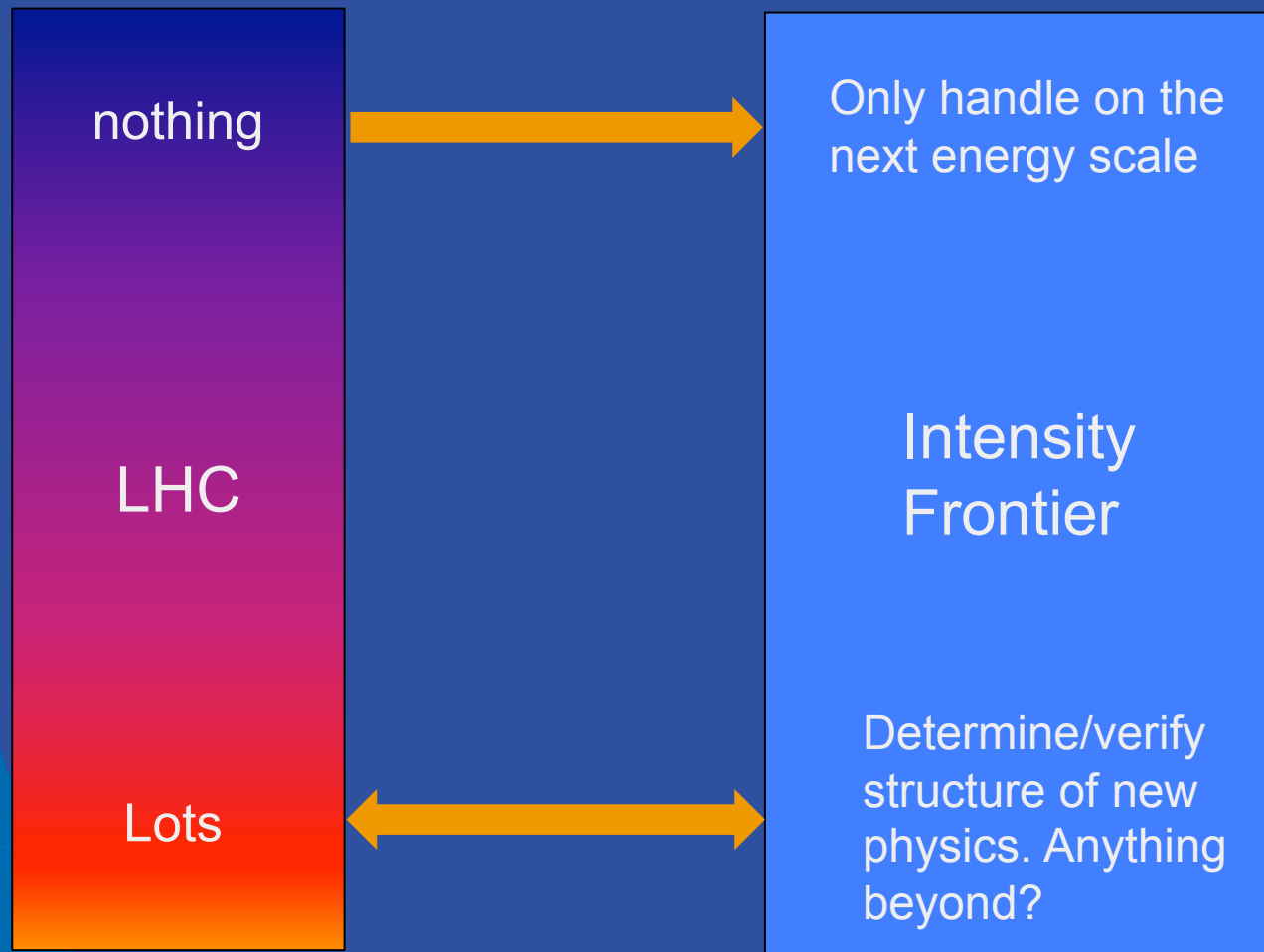


Color indicates degree of needed R&D (difficulty) and demonstration

The energy frontier



Interplay: LHC ↔ Intensity Frontier



Intensity is key for neutrinos

- Only weak interactions: very small cross sections >> hard to study
- Need large flux of particles and massive detectors
- Complementary to LHC: measure neutrino parameters (new symmetries?), neutrino masses, matter-antimatter symmetry violation and surprises.
- This route like the energy path depends of what we find in the current generation of experiments

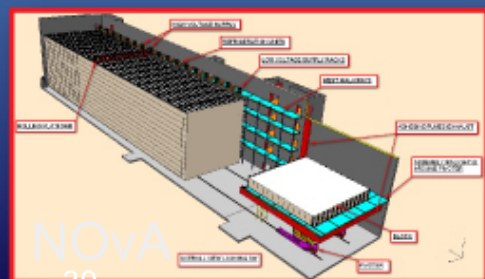
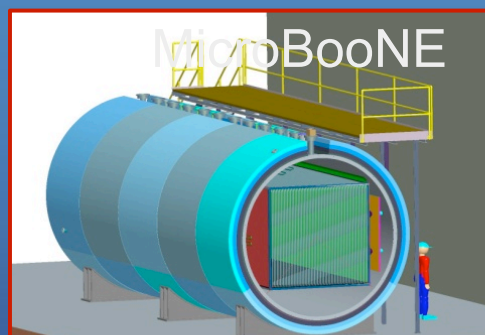
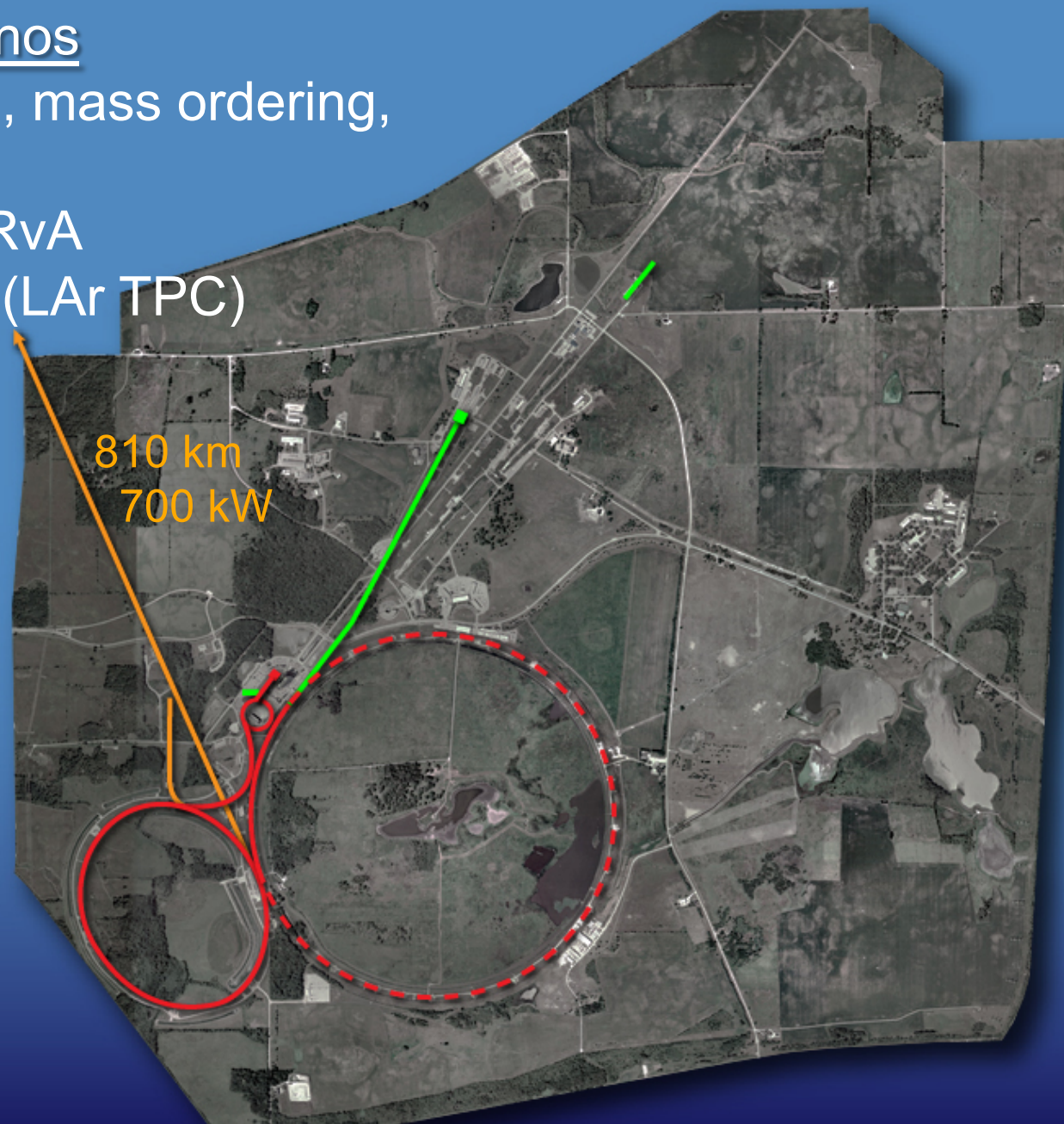
Neutrinos

NOvA (off-axis): θ_{13} , mass ordering,

....

MINERvA

MicroBooNE (LAr TPC)



Neutrinos

LBNE: FNAL \rightarrow DUSEL (+proton decay)

(DOE 1st stage approval)

Muons

Mu2e (DOE 1st stage approval)

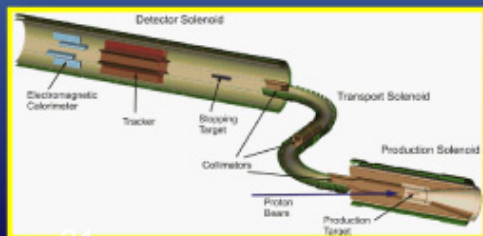
Muon g-2/EDM (to be reviewed by DOE)



700 kW
1300 km

Kaons

$K^+ \rightarrow \pi^+ \nu \bar{\nu}$
(proposed, but ...)



Project X

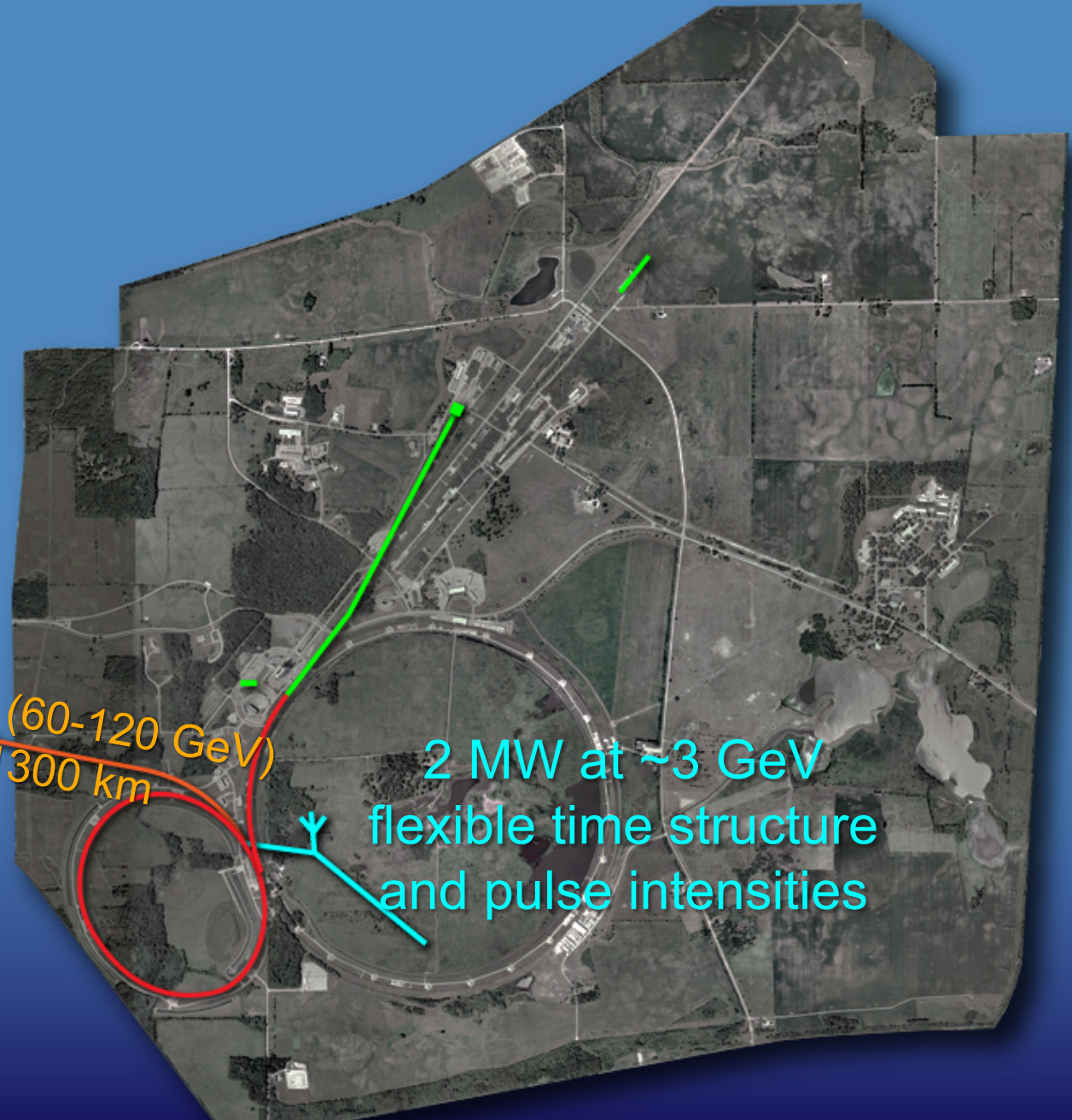
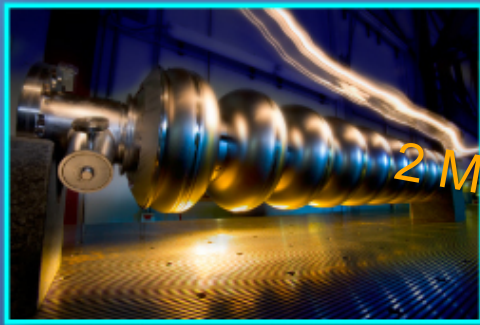
Neutrinos

Muons

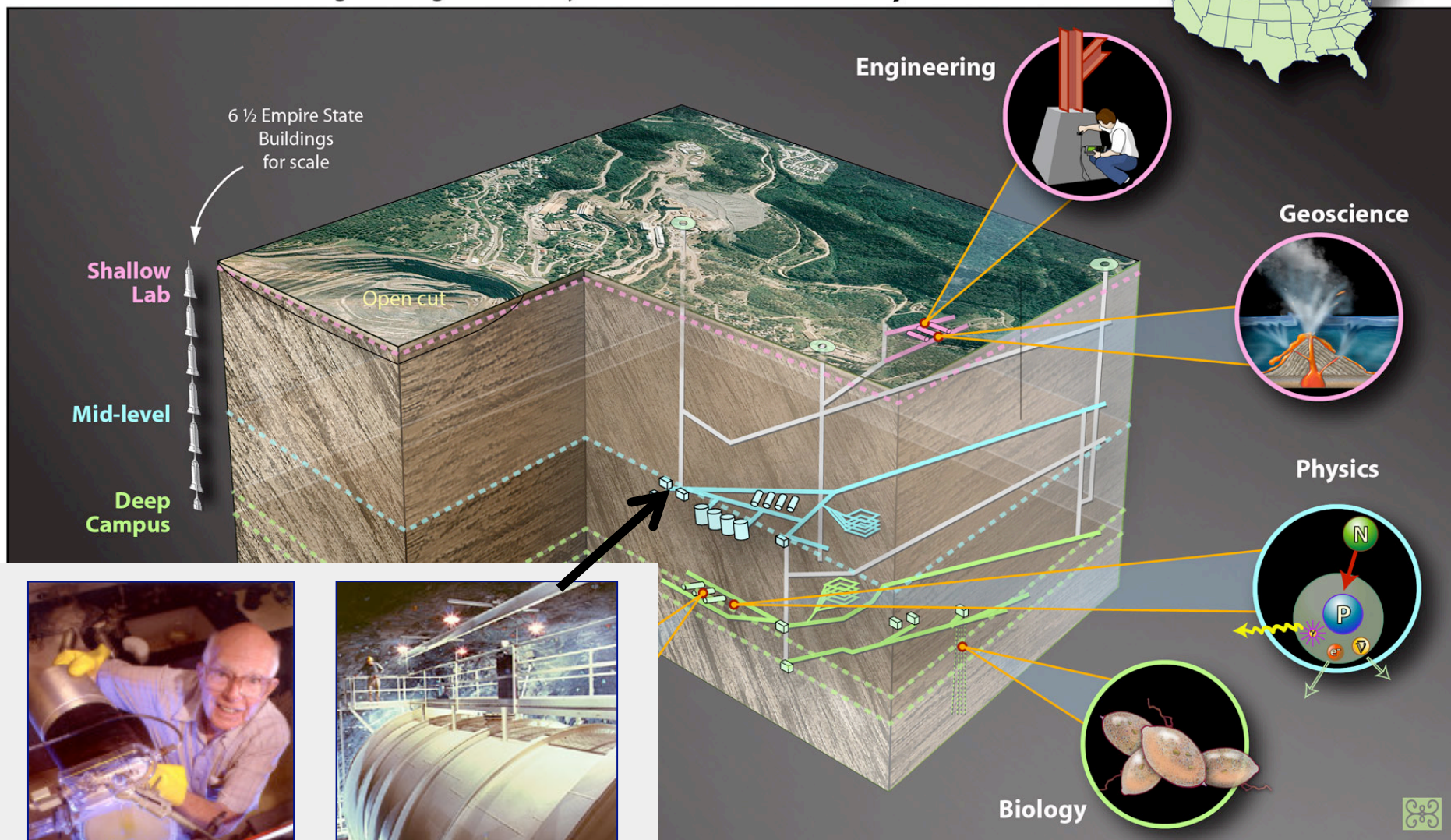
Kaons

Nuclei

“simultaneously”



DUSEL Deep Underground Science and Engineering Laboratory at Homestake, SD

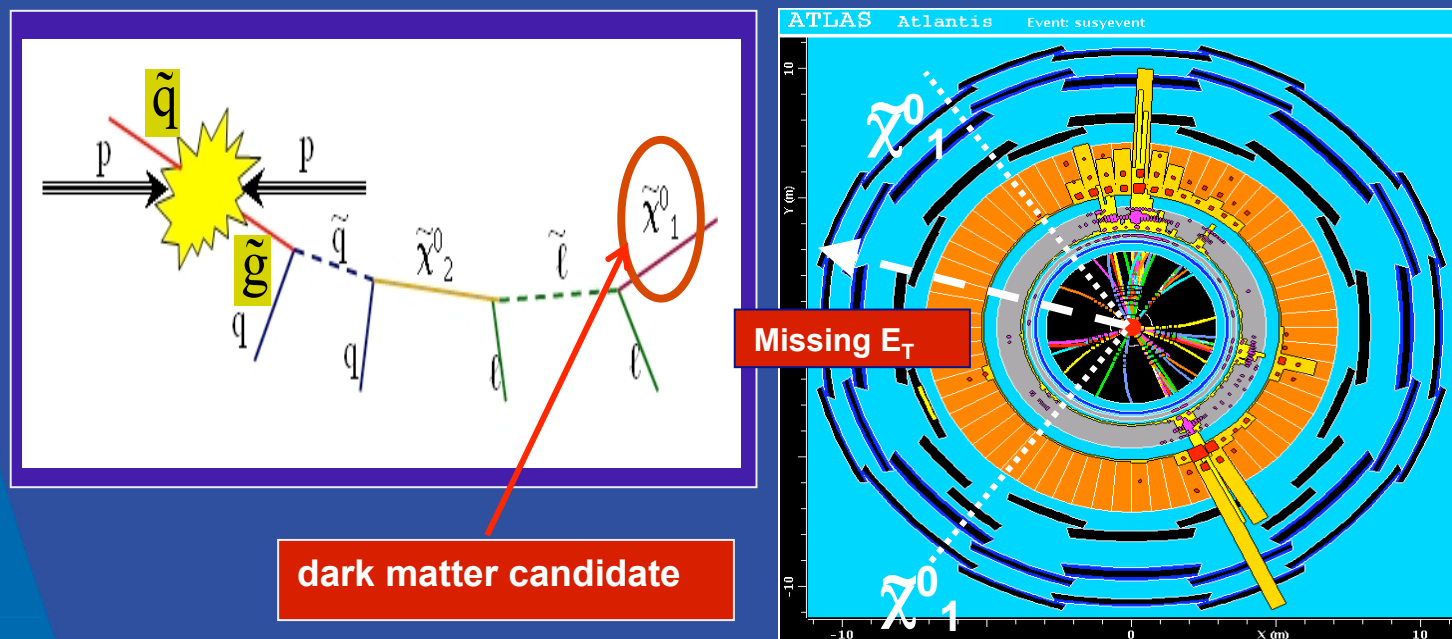


33 Ray Davis's Experiment

khaven Forum, May 28,2010

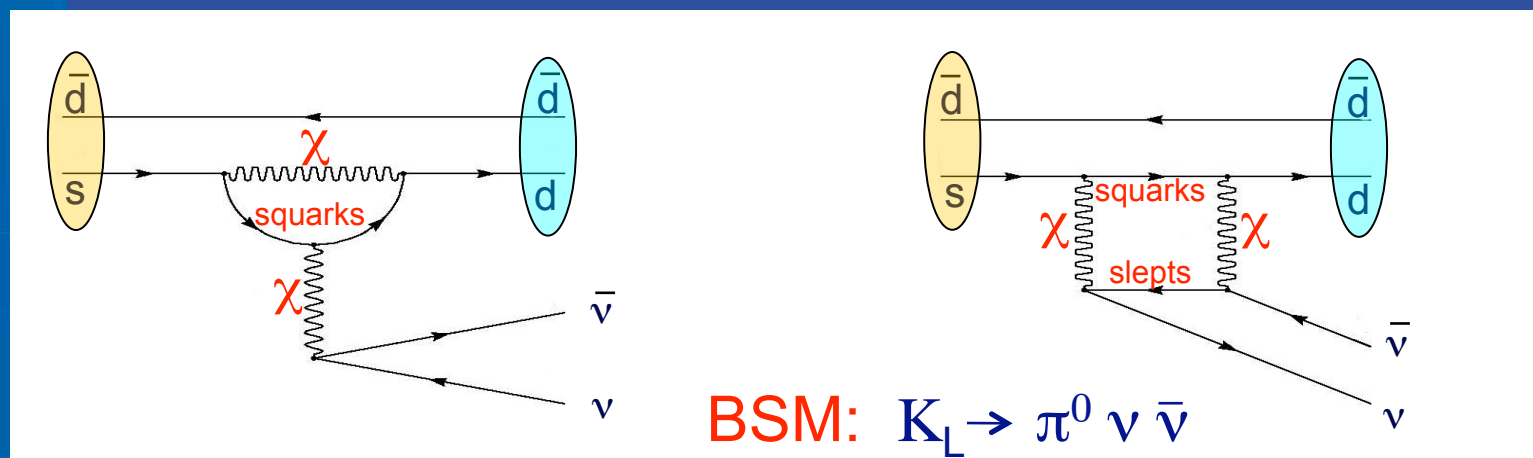
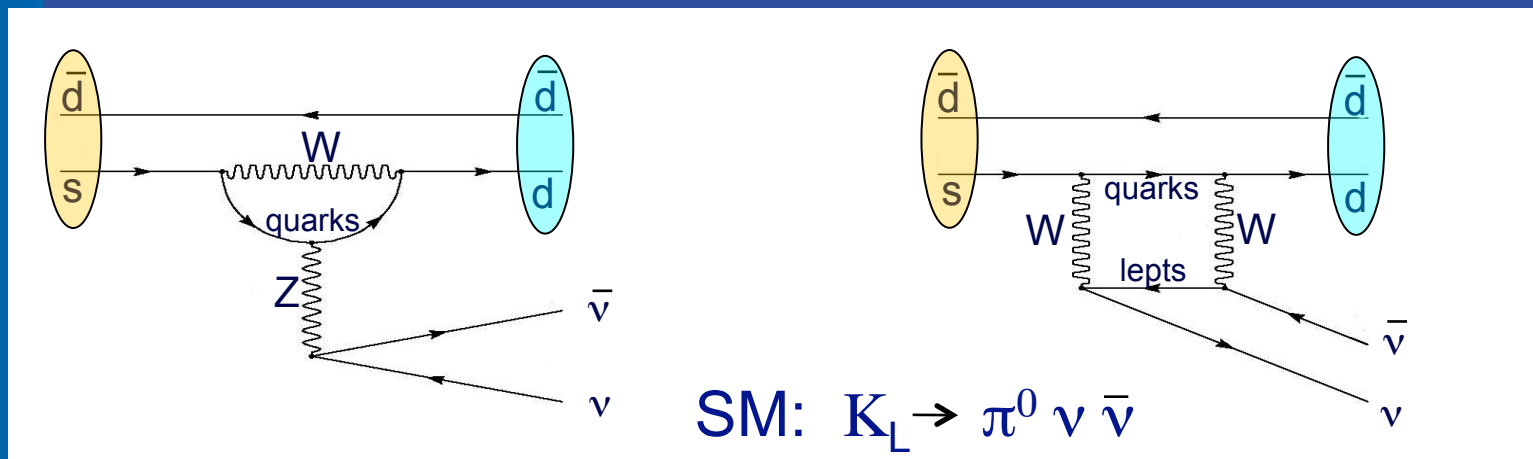
More LHC/Intensity frontier

- ATLAS/CMS discovers strongly coupled SUSY



- A host of new particles: fit roughly some masses, make assumption on couplings

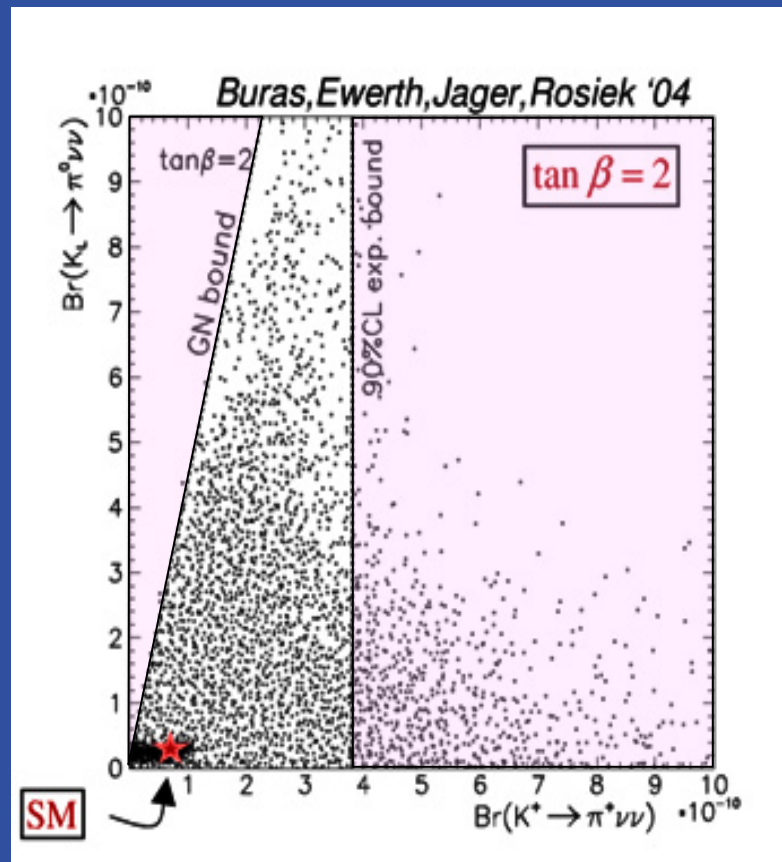
Large effects in kaon decay rates



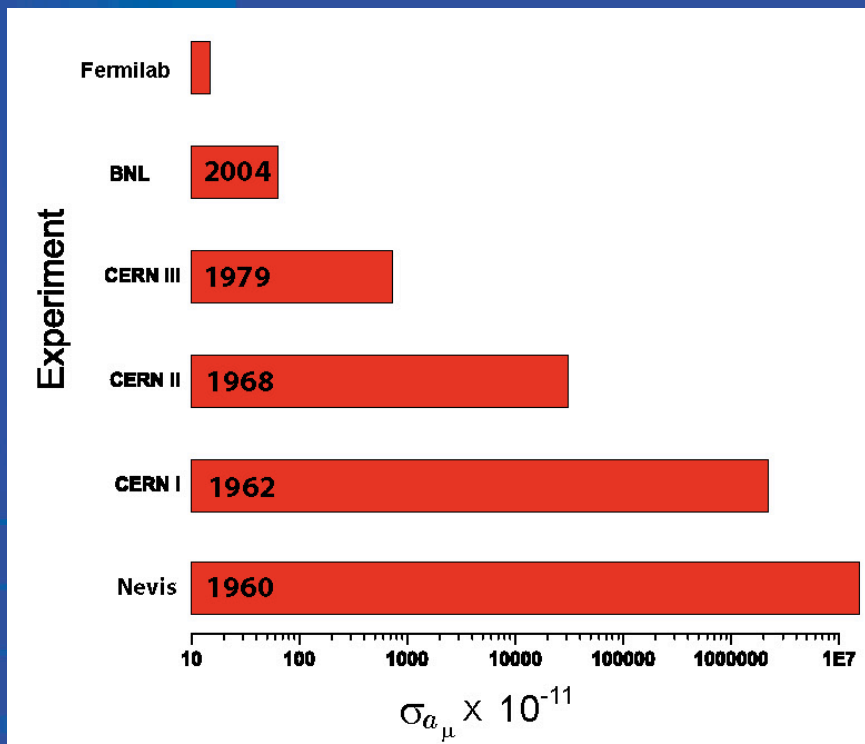
For particular classes of SUSY

Decay	Branching Ratio ($\times 10^{10}$)	
	Theory (SM)	Experiment
$K^+ \rightarrow \pi^+ \nu \bar{\nu} (\gamma)$	$0.85 \pm 0.07^{[1]}$	$1.73^{+1.15}_{-1.05}{}^{[2]}$
$K_L^0 \rightarrow \pi^0 \nu \bar{\nu}$	$0.28 \pm 0.04^{[3]}$	< 670 (90% CL) $^{[4]}$

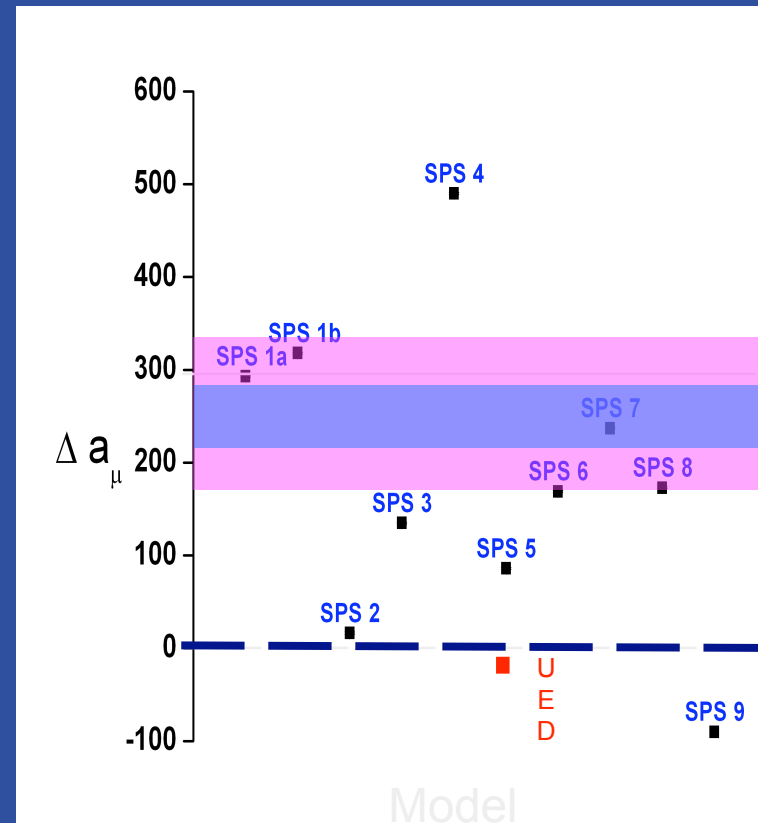
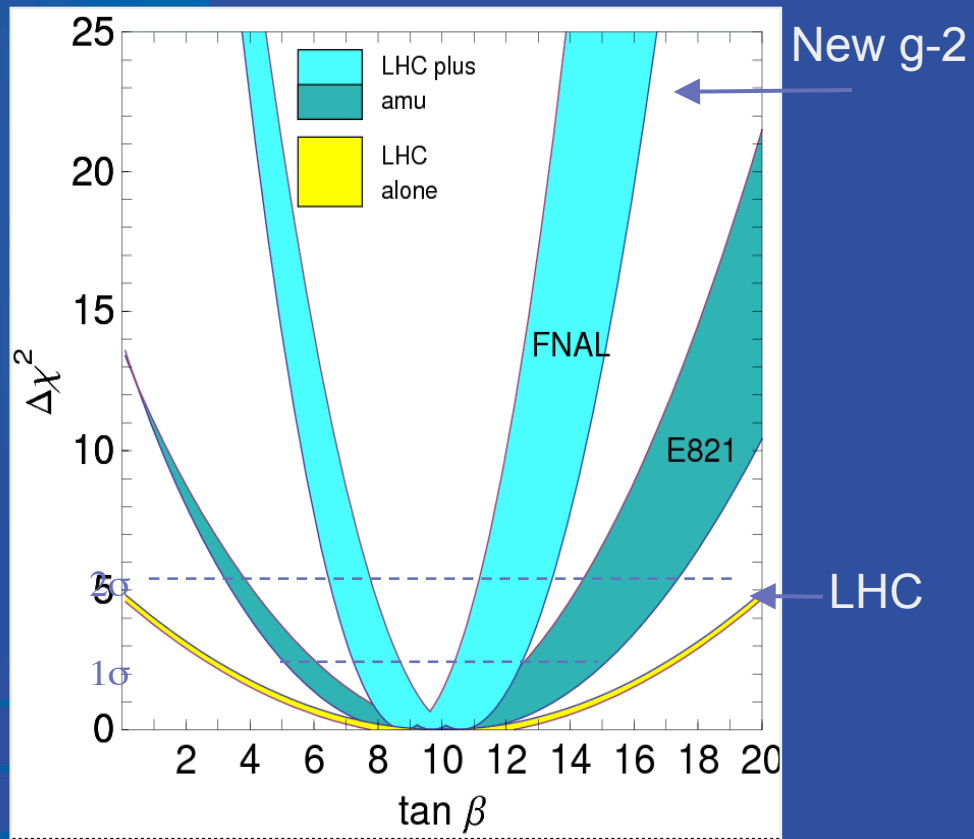
- Large effect on rare K decay modes highly suppressed with SM particles
- Much higher SM backgrounds in B and C decays
- (See also Neubert at BF2010)



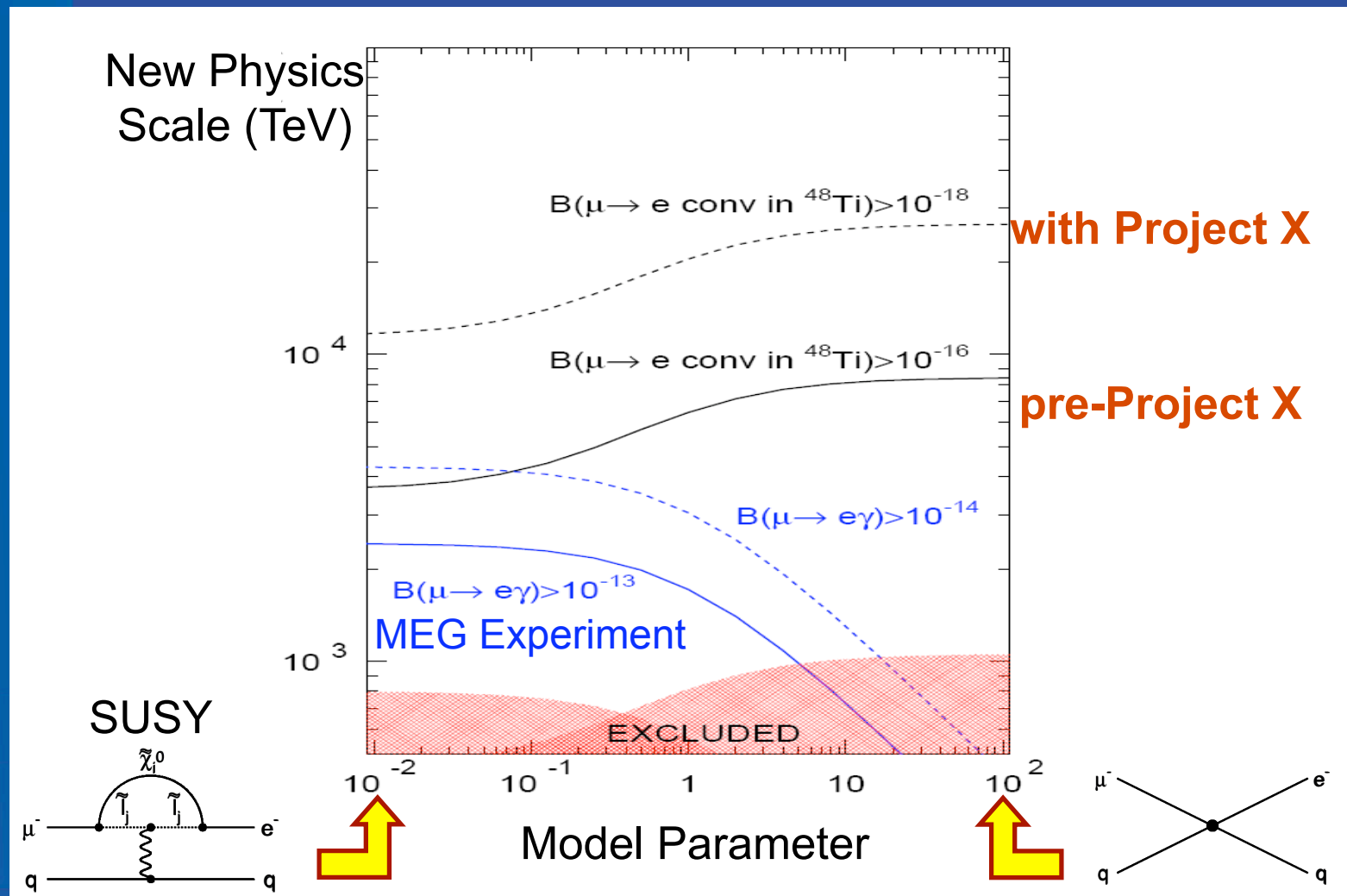
A new (g-2) to error of 0.14×10^{-11}



A new (g-2) to $0.14 \cdot 10^{-11}$



Mu2e can probe $10^3 - 10^4$ TeV

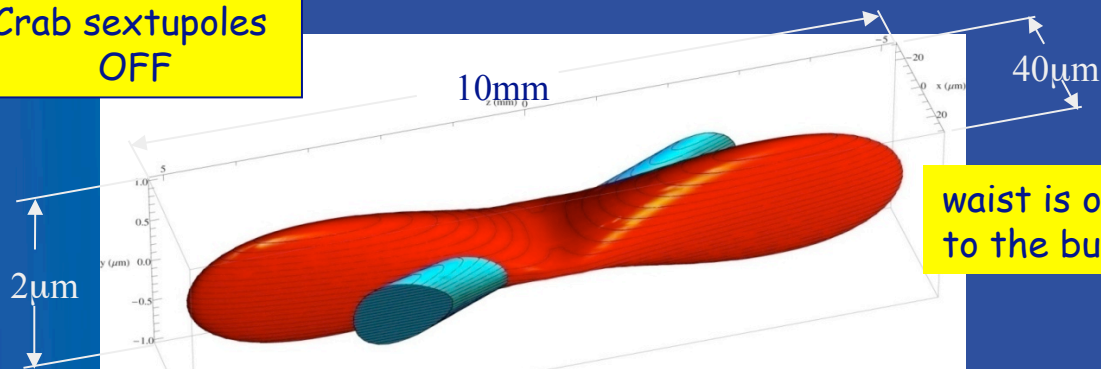


At the intensity frontier: Super B

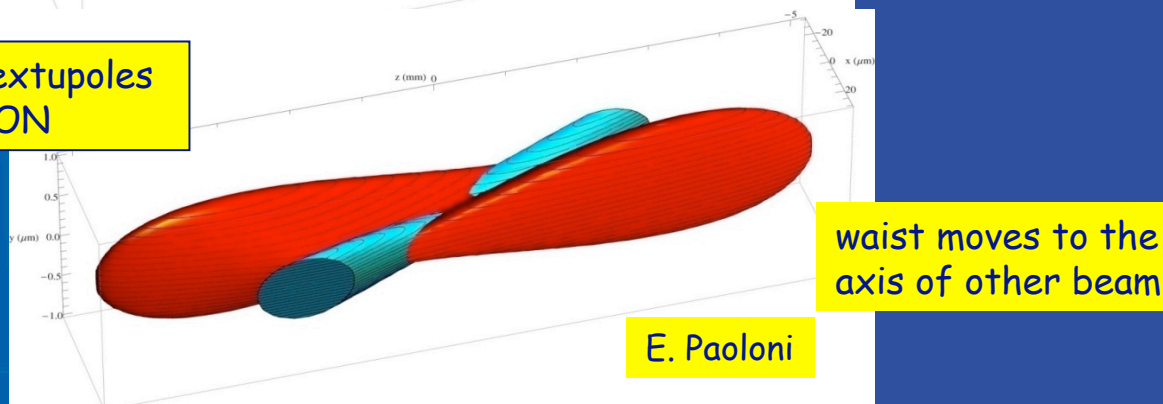
- One hundred times the luminosity of existing B-factories.
- Complementary program to LHC: flavor physics will manifest discoveries at LHC as well as higher mass scales
- Unlikely to be produced with present designs due to huge power loads: go to low emittances and waist focus. The main challenge is to maintain the low emittance. Two designs one in Japan and one in Italy

Super B: 4 GeV x 7 GeV

Crab sextupoles
OFF

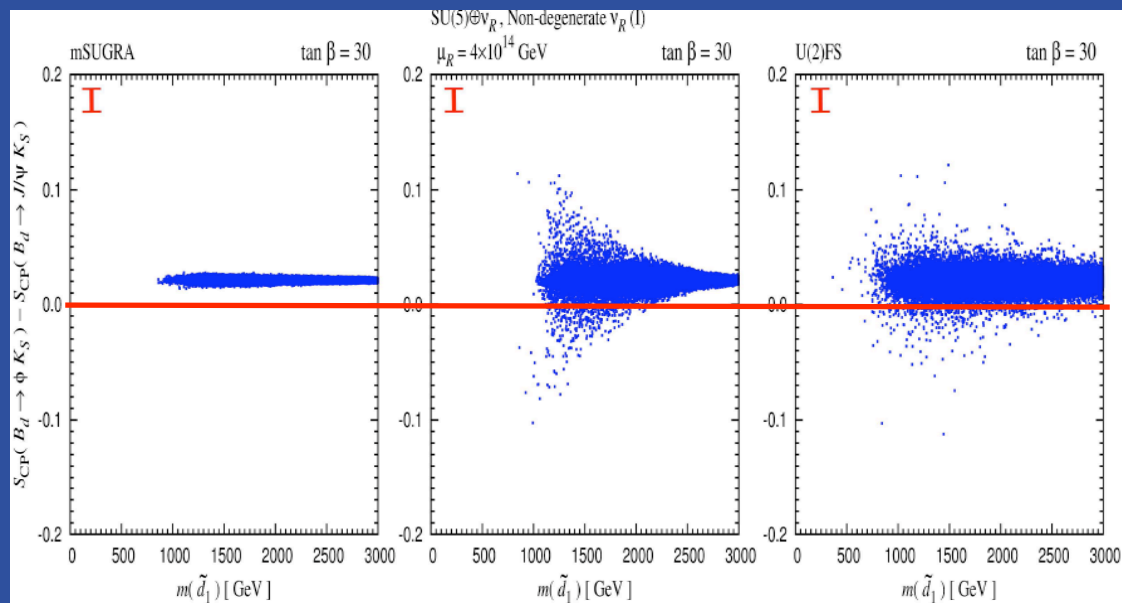
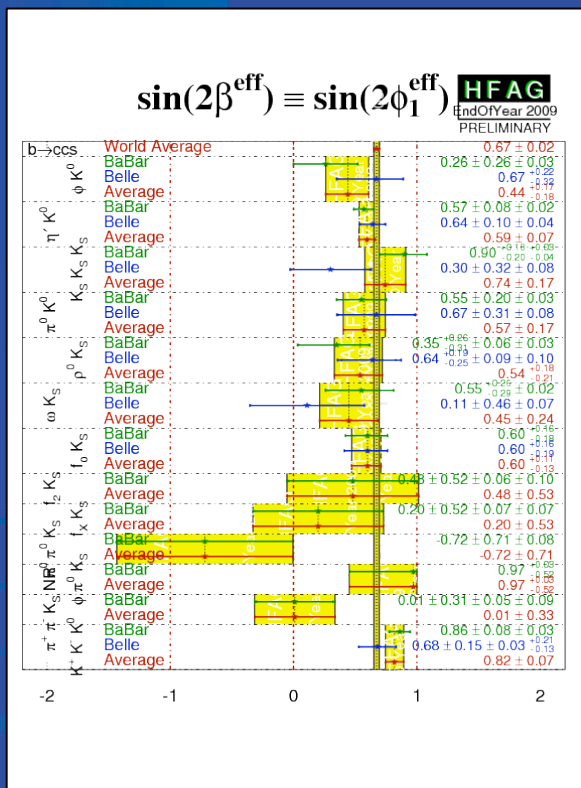


Crab sextupoles
ON



With crabbed waist, all particles from both beams collide in the minimum β_y region, producing a net gain in luminosity and a broad tune plane

New physics in CPV: $\sin 2\beta$



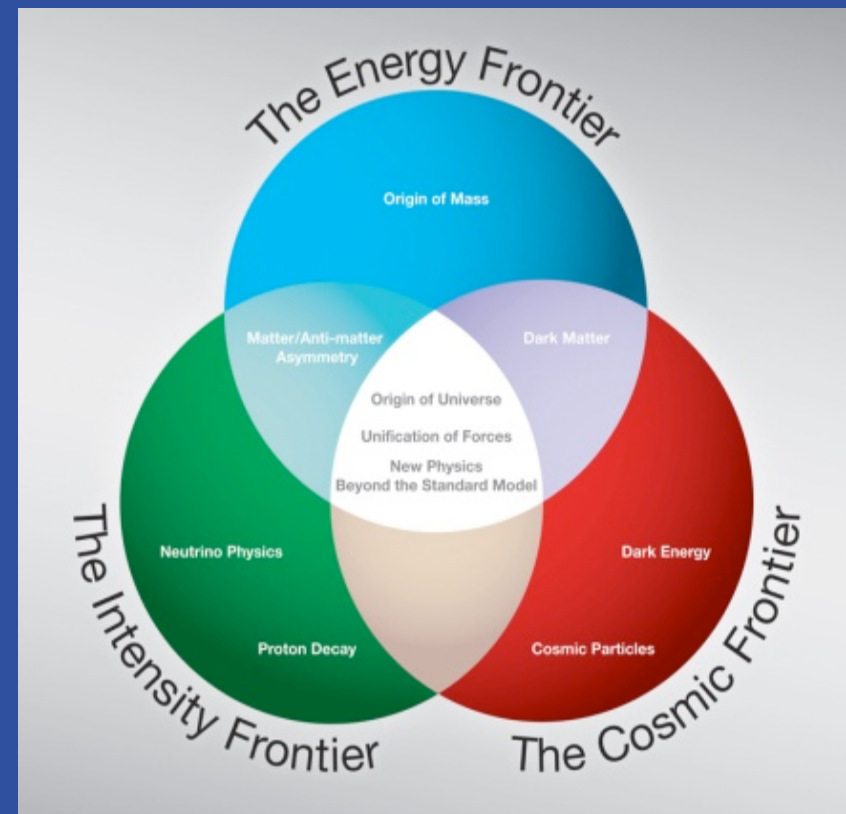
Many channels can show effects in the range $\Delta s \sim 0.01-0.04$

Observable	B Factories (2 ab^{-1})	Super B (75 ab^{-1})
$S(\phi K^0)$	0.13	0.02 (*)
$S(\eta' K^0)$	0.05	0.01 (*)
$S(K_s^0 K_s^0 K_s^0)$	0.15	0.02 (*)
$S(K_s^0 \pi^0)$	0.15	0.02 (*)
$S(\omega K_s^0)$	0.17	0.03 (*)
$S(f_0 K_s^0)$	0.12	0.02 (*)

(*) *theory limited*

Three main thrusts

- The energy frontier: produce particles at highest energy
- The intensity frontier: the most particles for neutrinos and rare decays
- Cosmic frontier: study phenomena in nature



In conclusion: the next decade.....

Chaos ~~or~~ illumination



Chaos and illumination